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Tropospheric jet response to Antarctic ozone depletion: An update with Chemistry-Climate Model Initiative (CCMI) models

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Abstract

The Southern Hemisphere (SH) zonal-mean circulation change in response to Antarctic ozone depletion is re-visited by examining a set of the latest model simulations archived for the Chemistry-Climate Model Initiative (CCMI) project. All models reasonably well reproduce Antarctic ozone depletion in the late 20th century. The related SH-summer circulation changes, such as a poleward intensification of westerly jet and a poleward expansion of the Hadley cell, are also well captured. All experiments exhibit quantitatively the same multi-model mean trend, irrespective of whether the ocean is coupled or prescribed. Results are also quantitatively similar to those derived from the Coupled Model Intercomparison Project phase 5 (CMIP5) high-top model simulations in which the stratospheric ozone is mostly prescribed with monthly- and zonally-averaged values. These results suggest that the ozone-hole-induced SH-summer circulation changes are robust across the models irrespective of the specific chemistry-atmosphere-ocean coupling.

1. Introduction

The Southern Hemisphere (SH) general circulation underwent distinct changes in the late 20th century. Among others, the westerly jet shifted poleward (Chen and Held 2007, Swart *et al* 2015), often represented by the positive trend of the southern annular mode index. The poleward edge of the Hadley cell also shifted poleward (Hu and Fu 2007, Davis and Rosenlof 2012, Garfinkel *et al* 2015), implying a widening of the Hadley cell. In response to these changes, SH surface climate variables such as surface air temperature and precipitation also changed significantly (Gillett *et al* 2006, Thompson *et al* 2011, Gonzalez *et al* 2014).

Simultaneous with these changes, Antarctic ozone concentrations sharply decreased due to the emission of ozone depleting substances (Solomon 1999). In an attempt to substantiate the causal link between the Antarctic ozone depletion and SH tropospheric and surface climate changes, multiple studies have performed climate model simulations with and without ozone depletion (e.g. Polvani *et al* 2011, McLandress *et al* 2011, Waugh *et al* 2015). A common feature of these studies is that a rapid decline of the austral-spring ozone concentrations in the lower stratosphere tends to force the austral-summer jet and Hadley cell to shift poleward. More importantly, these ozone-hole-induced circulation changes in austral summer are much stronger than the greenhouse-gas-induced ones. Although the detailed mechanism(s) remains to be determined, similar results are also seen in multi-model ensembles, e.g. the Coupled Model Inter-comparison Project (CMIP) phase 3 or 5 (Meehl *et al* 2007, Taylor *et al* 2012) and the Chemistry-Climate Model Validation activity 2 (CCMVal2; Eyring *et al* 2010), stressing that Antarctic ozone hole has played a predominant role in the *austral-summer* SH circulation changes in the late 20th century (Son *et al* 2009, Min and Son 2013, Gerber and Son 2014, Tao *et al* 2016, Choi *et al* 2018).

Only two studies do not conclude that ozone depletion dominated historical SH-summer circulation changes (Staten *et al* 2012, Quan *et al* 2014), and this can be partly traced back to the methodology used in their studies (Waugh *et al* 2015). It can be also influenced by the different sea surface temperature (SST) forcings (Staten *et al* 2012). However, the influence of SST variation on 20th century SH circulation changes is likely much weaker than the ozone-hole-induced ones (Waugh *et al* 2015). SST variations become important only after 2000 when ozone concentrations plateaued (Garfinkel *et al* 2015).

While the poleward intensification of the SH-summer jet in response to Antarctic ozone depletion is reasonably well simulated by most climate models, its magnitude differs substantially among models (e.g. Son *et al* 2009, Gerber and Son 2014). This inter-model spread could be caused by several, likely complementary factors. The most immediate factor

is the precise manner in which stratospheric ozone is imposed. While some models interactively simulate stratospheric chemistry and hence simulate an ozone hole with a three-dimensional structure that varies consistently with dynamical fields (e.g. CCMVal2 models), others simply prescribe stratospheric ozone using an off-line ozone dataset (e.g. CMIP3 and most CMIP5 models). Modeling studies have shown that the formers tend to simulate stronger tropospheric trends than the latter (Gillett *et al* 2009, Waugh *et al* 2009, Li *et al* 2016). This difference is caused not only by the realism of the ozone forcing but also by model biases in the simulation of the stratospheric polar vortex. Most CCMVal2 models, for example, suffer from a delayed break-up of the stratospheric polar vortex (Butchart *et al* 2011), and this bias can lead to an overestimate of the ozone-hole effect (Lin *et al* 2017). The ozone-hole-induced circulation change can be also sensitive to the temporal resolution and spatial structure of prescribed ozone: models prescribing daily and zonally asymmetric ozone often show stronger circulation changes than those forced by monthly and zonally-mean ozone (e.g. Crook *et al* 2008, Neely *et al* 2014).

However, the above sensitivities, which are mostly based on single model experiments with varying stratospheric ozone, do not explain differences between multi-model ensembles. For example, differences in the SH-summer circulation changes between CMIP3 simulations (where monthly- and zonally-averaged ozone is prescribed) and CCMVal2 simulations (where stratospheric ozone is fully interactive) are only minor (see figure 4 of Gerber *et al* 2012). Seviour *et al* (2017) also documented no systematic differences between model simulations with and without interactive ozone chemistry, and instead suggested that differences among simulations could reflect natural variability in the tropospheric circulation.

The SH-summer circulation changes could also be sensitive to surface boundary conditions. A poleward intensification of the jet can lead to cooler SST anomalies in high-latitudes but warm SST anomalies in mid-latitudes through the wind-driven meridional overturning circulation (Sigmond and Fyfe 2010, Thompson *et al* 2011). This SST change is then modified by a time-delayed deep ocean circulation change (Ferreira *et al* 2015, Seviour *et al* 2016). The net SST change differs substantially among the models (Ferreira *et al* 2015), introducing an uncertainty in the SH circulation change. Note that most CCMVal2 models were not configured with a coupled ocean (Morgenstern *et al* 2010), and hence the SST and sea ice concentration (SIC) did not evolve in a physically consistent manner with the overlying atmosphere.

The purpose of the present study is to re-assess the ozone-hole-induced tropospheric circulation changes by examining recent CCM simulations that were performed for the CCM Initiative (CCMI) project (Eyring *et al* 2013b, Morgenstern *et al* 2017). We address whether up-to-date CCMs, which have

Table 1. List of CCM1 models used in this study. Each model's acronym can be found in Morgenstern *et al* (2017). The model resolution is indicated in terms of horizontal resolution (longitude \times latitude) and the number of vertical levels. Models with only stratospheric chemistry are denoted with 'Strat', while those incorporating both stratospheric and tropospheric chemistry are denoted with 'Strat-Trop'. Models with relatively simple tropospheric chemistry are separately denoted with 'Strat-sTrop'. In the fourth column, 'Coupled' indicates the model in which the ocean is coupled in CCM1-C2 run.

Model	Resolution	Chemistry	CCM1-C2 ocean
ACCESS-CCM	3.75° \times 2.5° L60	Strat-Trop	Uncoupled
CCSRNIES-MIROC3.2	T42 L34	Strat	Uncoupled
CESM1-CAM4Chem	1.9° \times 2.5° L26	Strat-Trop	Coupled
CESM1-WACCM	1.9° \times 2.5° L66	Strat-Trop	Coupled
CMAM	T47 L71	Strat-Trop	Uncoupled
CNRM-CM5.3	T63 L60	Strat	Uncoupled
EMAC-L47MA	T42 L47	Strat-Trop	Coupled
EMAC-L90MA	T42 L90	Strat-Trop	Uncoupled
GEOSCCM	2° \times 2° L72	Strat-Trop	Uncoupled
HadGEM3-ES	1.875° \times 1.25° L85	Strat-Trop	Coupled
MRI-ESM1r1	T _L 159 L80	Strat-Trop	Coupled
NIWA-UKCA	3.75° \times 2.5° L60	Strat-Trop	Coupled
SOCOL3	T42 L39	Strat-sTrop	Uncoupled
UMSLIMCAT	3.75° \times 2.5° L64	Strat	Uncoupled
UMUKCA-UCAM	N48 L60	Strat-sTrop	Uncoupled

coupled ocean and more comprehensive chemistry, can represent a more realistic jet and its long-term trend compared with CCMVal2 simulations (Son *et al* 2010). Another purpose is to re-evaluate the importance of interactive ozone chemistry and a coupled ocean. This issue was recently addressed by Seviour *et al* (2017), who performed time-slice experiments with varying stratospheric ozone forcing with and without a coupled ocean, but is extended in this study to multi-model transient simulations. For this purpose, the CCM1 model simulations with and without a coupled ocean are directly compared. The multi-model mean trend of the CCM1 simulations is also compared with that of the CMIP5 simulations.

Here it should be stated that the models analyzed in this study are not solely driven by ozone depletion. Other forcings, such as increasing greenhouse gas concentrations and anthropogenic aerosol loadings, are also included. But, based on previous studies (e.g. Polvani *et al* 2011, Waugh *et al* 2015), it is assumed that the SH-summer circulation changes in the late 20th century are mostly driven by Antarctic ozone depletion.

2. CCM1 and CMIP5 datasets

The CCM1 was jointly launched by the International Global Atmospheric Chemistry (IGAC) and the Stratosphere-troposphere Processes And their Role in Climate (SPARC) to better understand chemistry-climate interactions in the recent past and future climate (Eyring *et al* 2013b). This modeling effort is an extension of CCMVal2 (Eyring *et al* 2010), but utilizes up-to-date CCMs. The CCM1 models used in this study are listed in table 1. All models that provide the reference simulations of the recent past and future climate are considered. Models with missing data or low resolution (coarser than T42 resolution) are excluded. As briefly described in table 1, tropospheric chemistry, in addition to stratospheric chemistry, is

fully interactive in most models (Morgenstern *et al* 2017). This differs from most of the CCMVal2 models in which only stratospheric chemistry is interactive (Morgenstern *et al* 2010). More importantly, six CCM1 models (i.e. CESM1-CAM4Chem, CESM1-WACCM, EMAC-L47MA, HadGEM3-ES, MRI-ESM1r1, and NIWA-UKCA) are integrated with a coupled ocean (Morgenstern *et al* 2017), enabling us to evaluate the role of chemistry-atmosphere-ocean coupling in SH climate change.

Two sets of CCM1 simulations, i.e. REF-C1 and REF-C2, are investigated in this study (Eyring *et al* 2013b, Morgenstern *et al* 2017). The CCM1 REF-C1 (hereafter referred to as CCM1-C1) experiment is a historical simulation, forced by observed SST/SIC. In contrast, the CCM1 REF-C2 (hereafter CCM1-C2) experiment covers not only historical period but also future climate. This experiment is conducted either with a coupled ocean or with modeled SST/SIC taken from coupled model simulations (e.g. CMIP5). The one-to-one comparison of these two experiments thus allows us to quantify the importance of surface boundary conditions.

To identify the importance of interactive chemistry, CCM1 simulations are also compared with CMIP5 historical simulations. Only the high-top CMIP5 models, which have a model top at 1 hPa or higher, are considered in this study (table 2). Most of them are forced by the SPARC ozone data or its modified version (Eyring *et al* 2013a). However, several models have fully interactive ozone chemistry and can be considered as CCMs (four models in table 2). In fact, two of them, i.e. CESM1-WACCM and MRI-ESM1, participated in the CCM1 project. Not surprisingly, these models have quantitatively different ozone evolution from the SPARC ozone data. However, for simplicity, multi-model mean trends are constructed by averaging all CMIP5 high-top model simulations without consideration of the details of ozone chemistry. A comparison between the CMIP5 models with and without interactive chemistry is only briefly discussed.

Table 2. List of CMIP5 models used in this study. Only high-top models, that have a model top at 1 hPa and higher, are used. Models prescribing ozone depletion are denoted with ‘Prescribed’, while those incorporating semi-offline chemistry or fully interactive ozone chemistry are denoted with ‘Semi-offline’ or ‘Strat-Trop’, respectively. Note that, unlike CCM1 models, all CMIP5 models are coupled with an ocean.

Model	Resolution	Chemistry
CESM1-WACCM	$1.9^\circ \times 2.5^\circ$ L66	Strat-Trop
CMCC-CMS	T63 L95	Prescribed
GFDL-CM3	C48 L48	Strat-Trop
HadGEM2-CC	N96 L38	Prescribed
IPSL-CM5A-LR	$1.875^\circ \times 3.75^\circ$ L39	Semi-Offline
IPSL-CM5A-MR	$1.25^\circ \times 2.5^\circ$ L39	Semi-Offline
IPSL-CM5B-LR	$1.875^\circ \times 3.75^\circ$ L39	Semi-Offline
MIROC4h	T213 L56	Prescribed
MIROC-ESM	T42 L80	Prescribed
MIROC-ESM-CHEM	T42 L80	Strat-Trop
MPI-ESM-LR	T63 L47	Prescribed
MPI-ESM-MR	T63 L95	Prescribed
MPI-ESM-P	T63 L47	Prescribed
MRI-CGCM3	T159 L48	Prescribed
MRI-ESM1	T _L 159 L48	Strat-Trop

All analyses are conducted with the first ensemble member of each model. All model output is interpolated onto a common resolution of 2.5° latitude by 2.5° longitude and 31 pressure levels before computing linear trends. Although model output is available even in the 2000s, only the period of 1960–2000 when Antarctic ozone depletion is well defined is considered. Since the same analysis period has been used in the literature (Son *et al* 2010, Eyring *et al* 2013a, Gerber and Son 2014, Garfinkel *et al* 2015), this allows a direct comparison with previous studies.

3. Results

Figure 1 presents the evolution of September–November (SON) total column ozone (TCO) anomaly, area-weighted from 60°S to the pole, for each model. Here, the anomaly is defined as the deviation from the 1980–2000 climatology of each model. All models, i.e. CCM1-CI, CCM1-C2, and CMIP5 models, reasonably well reproduce a reduction in TCO as has been observed (Bodeker *et al* 2005, Van der *et al* 2015). The spatial distribution of the monthly-mean TCO and its seasonal cycle are also reasonably well captured (not shown). A comparison between figures 1(a) and (b) further reveals that, for each model, CCM1-C1 and CCM1-C2 simulations have a quantitatively similar TCO evolution (compare the same color on each panel). Since the two experiments differ mainly in surface boundary conditions (e.g. SST and SIC), this result may suggest that stratospheric ozone chemistry and transport is only weakly sensitive to the details of surface boundary conditions.

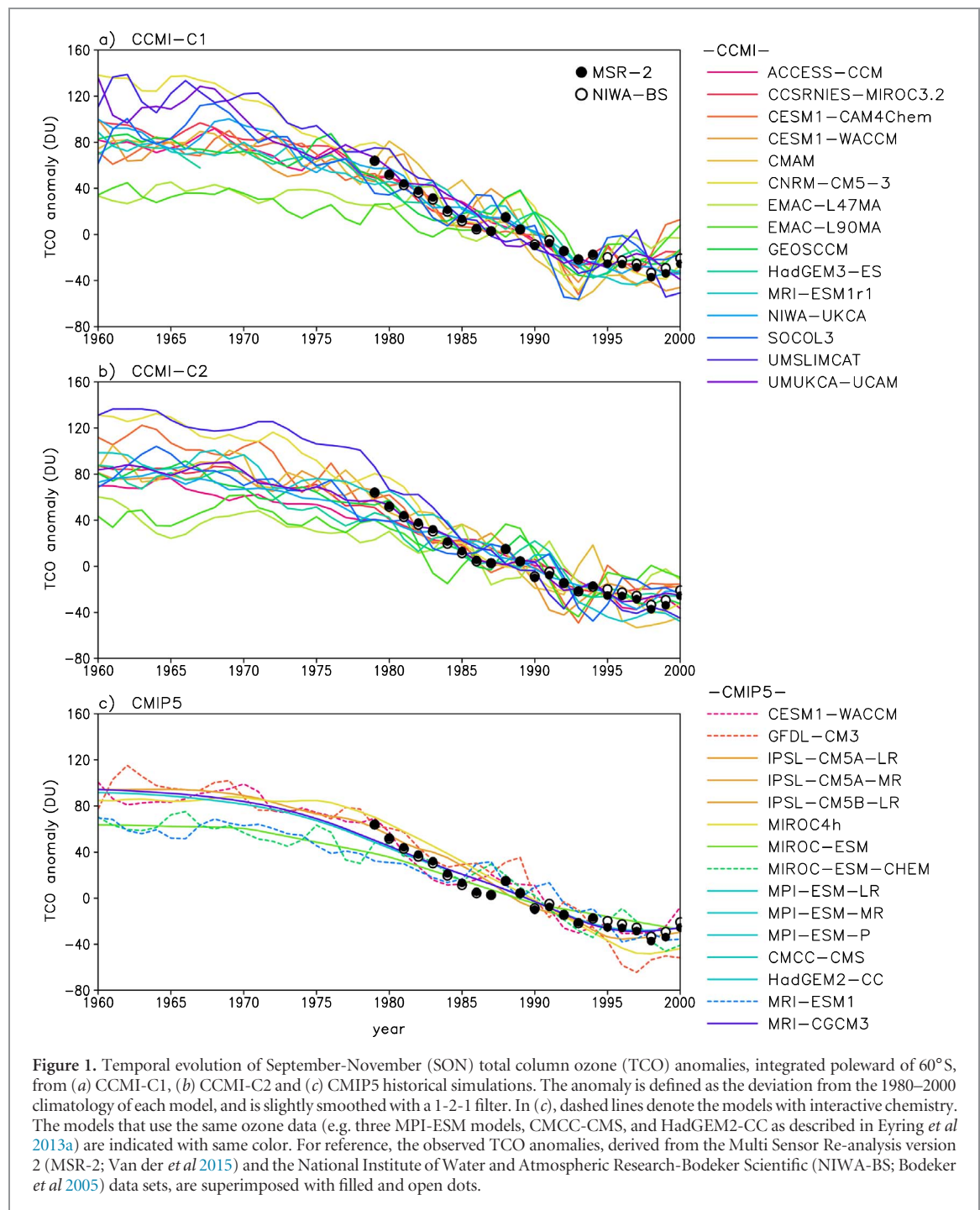
A pronounced inter-model spread, however, is evident especially in the 1960s and 1970s (figures 1(a) and (b)). This divergence among the models is similar to that seen in the CCMVal2 models (Austin *et al* 2010), and indicates that CCMs still have a large uncertainty

in their simulation of Antarctic ozone. Unlike CCM1 models, CMIP5 models show a quantitatively similar TCO evolution among the models (figure 1(c)). This is anticipated because most CMIP5 models are forced by the SPARC ozone data or its modified version. But, the CMIP5 models with interactive ozone chemistry also show a similar TCO evolution (see the dashed lines in figure 1(c)).

The vertical structure of polar ozone trends is illustrated in the first column of figure 2. The ozone depletion in the CCM1 simulations is maximum at ~ 50 hPa in October. This is quantitatively similar to the one derived from the SPARC ozone data (e.g. figure 2(c)). The subtle differences between the experiments, such as a stronger upper-stratospheric ozone depletion in CCM1 runs than in CMIP5 runs, are mostly insignificant.

The temperature response to the ozone depletion and the related stratospheric circulation change is shown in the middle column of figure 2. All experiments show significant cooling trend, centered at ~ 70 hPa, in November. This cooling trend extends from the middle stratosphere to the lower stratosphere with a maximum cooling at 20 hPa in October but at 200 hPa in December. Due to the thermal wind balance, a strong cooling in late spring is accompanied by a strengthening of the polar vortex (third column of figure 2). However this acceleration, which is a maximum in November, is not confined within the stratosphere but extends down into the troposphere. A statistically significant trend in the troposphere is particularly evident in December.

The temporal and vertical structure of polar ozone, temperature and mid-latitude wind trends, shown in figure 2, is remarkably similar to that of CCMVal2 simulations (figures 3(a)–(c) of Son *et al* 2010), indicating no major difference in multi-model mean trends between the CCMVal2 models and their updated versions. Such a similarity is also found in the December–February (DJF) zonal-mean zonal wind trends (figures 3(b)–(d)). For reference, the trend derived from the Japanese 55 year Reanalysis (JRA-55; Ebata *et al* 2011) is also displayed in figure 3(a). This data is chosen simply because it is the latest reanalysis dataset covering the analyzed period. All experiments show quantitatively a similar poleward intensification of westerly jet that resembles the reanalysis trend and the CMIP3/CCMVal2 trends (e.g. figure 4 of Gerber *et al* 2012). Note that CMIP5 models show a somewhat weak polar vortex change compared with CCM1 models (figure 3(d)). This underestimation is mainly due to the models prescribing ozone depletion (figure 3(f)). Models with interactive and semi-offline chemistry (seven models listed in table 2) show essentially the same jet trend as CCM1 models (figure 3(e)). However, regardless of polar vortex changes, two groups of CMIP5 models show quantitatively similar tropospheric circulation changes (compare figures 3(e) and (f)). This result is consistent with Eyring *et al* (2013a)

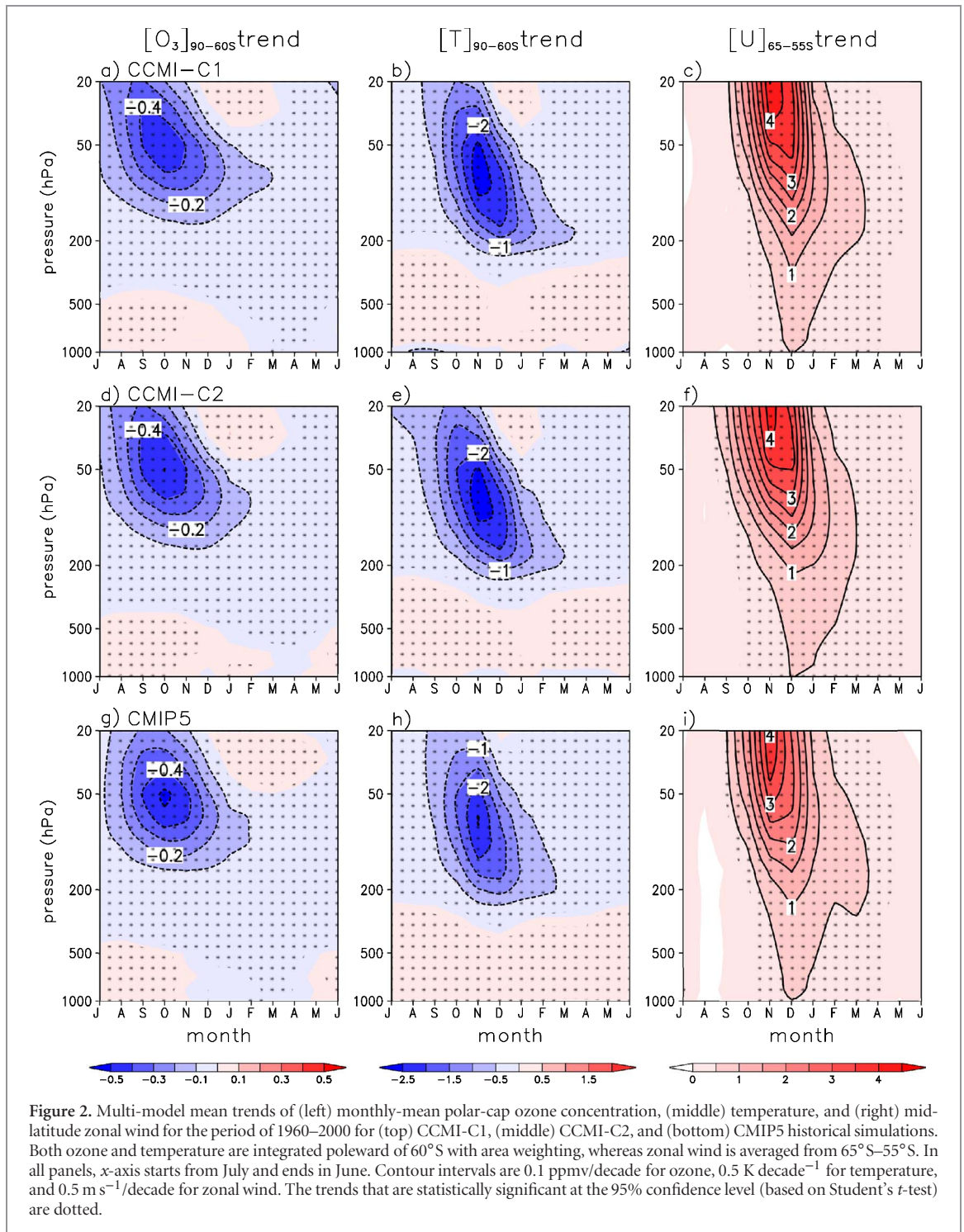


who documented that CMIP5 models with interactive chemistry have larger inter-model spread and are not well separated from those without interactive chemistry (see their figures 10 and 11).

These results clearly suggest that the ozone-hole-induced tropospheric changes are not strongly sensitive to the coupled ocean and interactive chemistry when the multi-model mean trends are considered. This conclusion is supported by 850 hPa zonal wind trends (figure 3(g)). Their latitudinal distributions are almost identical among the model ensembles. The CCMI–C1 models with observed SST/SIC and the same models with a coupled ocean (CCMI–C2) are separately compared in figure 3(h) (blue and green solid lines; see

table 1 for the six models with a coupled ocean). Likewise, CCMI–C1 models with observed SST/SIC and the same models with prescribed SST/SIC from the coupled models are compared (blue and green dashed lines). Each group again shows quantitatively the same multi-model mean trend, confirming the above conclusion.

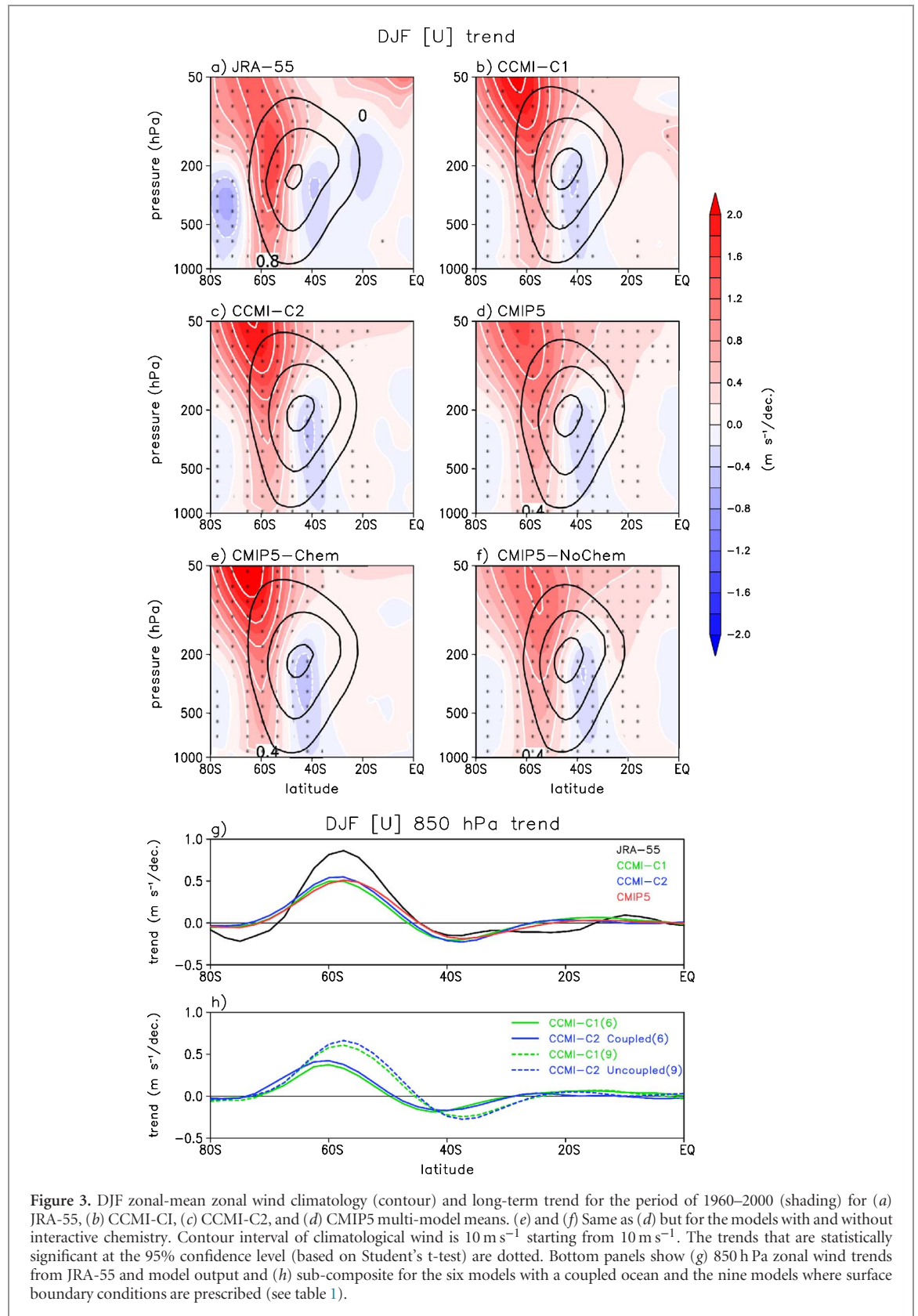
All analyses so far are based on multi-model mean trends. Individual models, however, exhibit significantly different trends. For instance, the six CCMI–C2 runs with a coupled ocean and the other nine in figure 3(h) show non-negligible differences (compare blue solid and dashed lines). These differences can be partly traced back to different magnitudes of ozone



depletion rather than different surface boundary conditions. Figure 4(a) shows the inter-model spread of DJF jet-latitude trends against polar-stratospheric ozone trends. Following previous studies (e.g. Son *et al* 2009), the jet latitude is determined as the latitude where a cubic fitted 850 hPa zonal-mean zonal wind, around its maximum grid point and the two points either side, has the maximum. The lower-stratospheric ozone trends are quantified by the October–January (ONDJ) ozone trend at 100 hPa, area-weighted from 60°S to the pole, based on figure 2(a).

The CCMI simulations show a wide range of jet-latitude trends from $\sim 1.2^\circ$ /decade poleward shift to $\sim 0.3^\circ$ /decade equatorward shift (figure 4(a)). However, only half of them are statistically significant (light shaded symbols), indicating a large uncertainty in the simulated jet trends. Not surprisingly, the inter-model spread of the jet-latitude trends is highly correlated with that of the polar-stratospheric ozone trends. Their correlation is about 0.65 for both CCMI-C1 and CCMI-C2 simulations.

Figure 4(c) presents the one-to-one comparison between CCMI-C1 and CCMI-C2 simulations. The



models with a coupled ocean in CCMI-C2 runs and those with observed SST/SIC in CCMI-C1 runs show no systematic differences (see purple bars; see also figure 3(h)). While three models show a weaker poleward jet shift when coupled to an ocean, the other three models show a stronger poleward jet shift. Moreover, none of these differences are statistically significant

(see open bars). The CCMI-C2 runs with modeled SST/SIC and CCMI-C1 runs with observed SST/SIC also show no systematic differences (cyan bars). Although two models (i.e. EMAC-L90MA and SOCOL3) show statistically significant differences, they are opposite in sign. Although not presented, essentially the same results are found when the jet-latitude

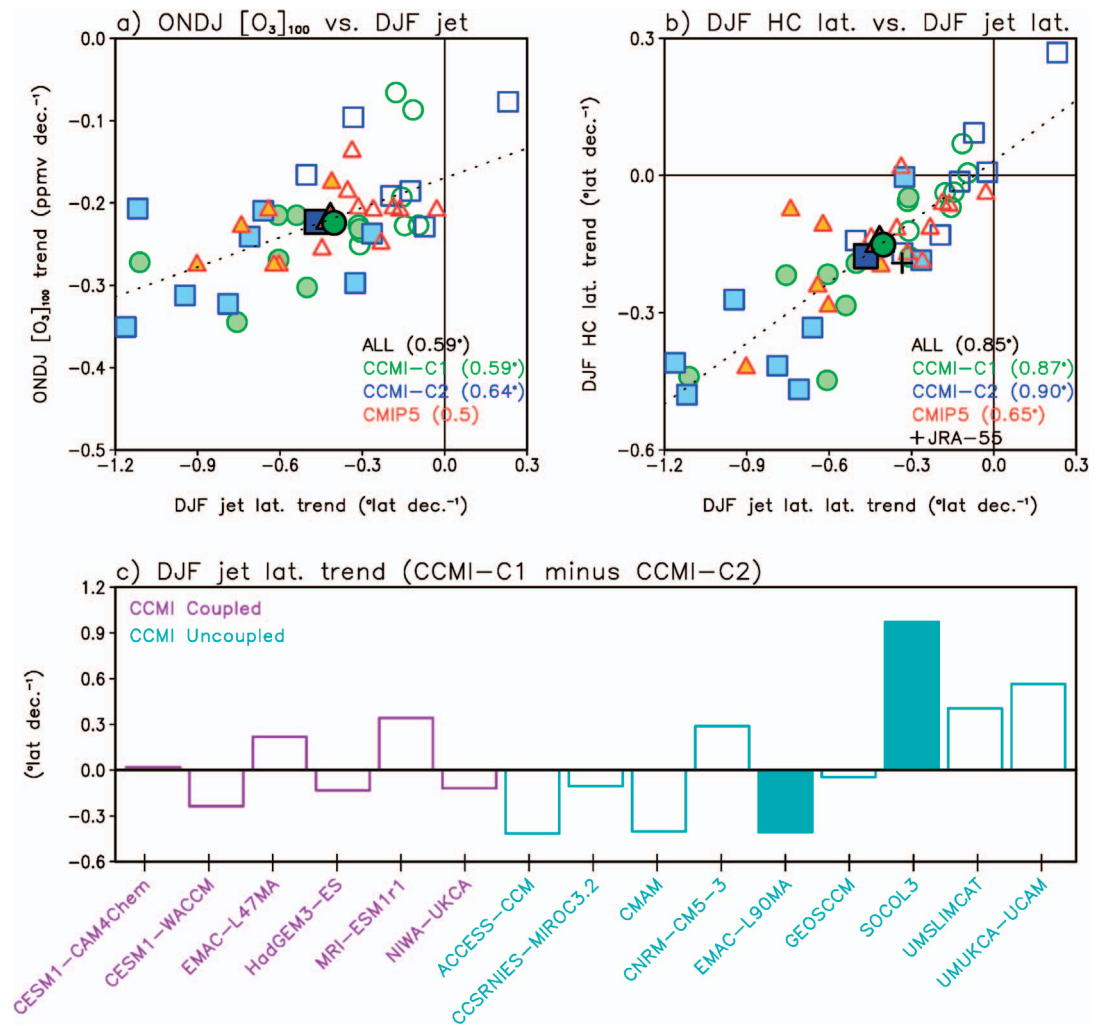


Figure 4. The relationship of DJF jet-latitude trends (a) to ONDJ polar-stratospheric ozone trends and (b) to DJF Hadley-cell edge trends for the period of 1960–2000. The jet latitude is determined with 850 h Pa zonal-mean zonal wind, while polar-stratospheric ozone is defined by 100 h Pa ozone area-weighted from 60°S to the pole. Models that show statistically significant jet-latitude trends are denoted with light colored symbols, while those with insignificant trends are denoted with open symbols. The cross and filled symbols indicate the trends derived from JRA-55 and multi-model means. Correlation coefficient for each experiment is indicated in the parenthesis, following the experiment name, with an asterisk if the value is statistically significant at the 95 % confidence level based on Student's *t*-test. Bar graph in (c) shows the jet latitude trend difference between CCMI-C1 and CCMI-C2 simulations. When the difference is statistically significant at the 95% confidence level, it is filled.

trends are normalized by the polar-stratospheric temperature trends. This confirms that the ozone-hole-induced SH-summer circulation changes are not strongly sensitive to the details of the surface boundary conditions (Seviour *et al* 2017).

All analyses are repeated for the Hadley cell edge. Here, the poleward edge of the Hadley cell is determined by the zero-crossing latitude of 500 h Pa mass stream function in the SH subtropics. During the austral summer, its trends are highly correlated with the jet-latitude trends (figure 4(b)). Their correlations across CCMI-C1, CCMI-C2, and CMIP5 runs are 0.87, 0.90, and 0.65, respectively. Consistent with previous studies (e.g. Son *et al* 2009), their ratio is close to 1-to-2 (see dashed line). Given this linear relationship, it is not surprising to find that overall results of the Hadley-cell changes are quite similar to the jet changes described above (not shown).

Previous ensembles of CCMs, such as CCMVal1, CCMVal2, and CMIP5 models with interactive chemistry, have suffered from biases in their mean state. Most CCMVal2 models, for instance, exhibit equatorward biases in the position of the climatological jet (see figure 10(b) of Son *et al* 2010). In terms of the multi-model mean, this bias is somewhat reduced in the CCMI simulations (see dark filled symbols in figure 5). However, compared to CCMVal2 models, the inter-model spread becomes larger with almost half of the models showing a poleward-biased climatological jet.

This bias has been related to different circulation responses to an identical forcing. Specifically, it has been proposed that models with an equatorward-biased jet tend to have a stronger jet response to the external forcing (Son *et al* 2010, Kidston and Gerber 2010). This argument, however, was questioned by recent studies (Simpson and Polvani 2016, Seviour *et al* 2017).

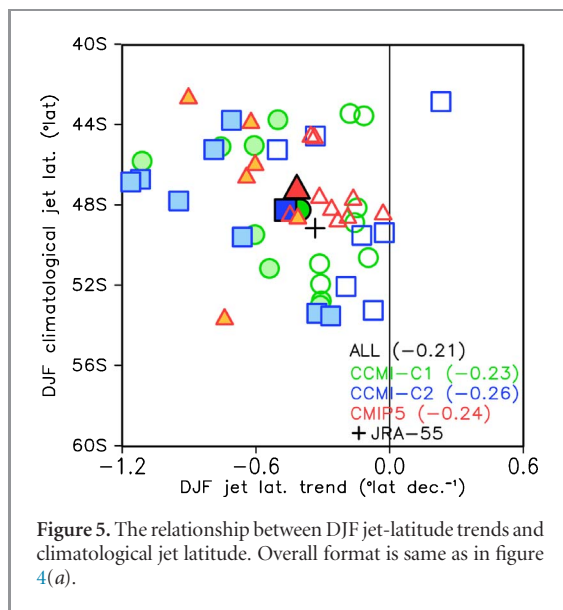


Figure 5. The relationship between DJF jet-latitude trends and climatological jet latitude. Overall format is same as in figure 4(a).

Figure 5 displays the relationship between jet-latitude trends and climatological jet locations. Although there is a hint of a linear relationship (i.e. models with an equatorward-jet bias tend to have a stronger jet trend), all three model ensembles show statistically insignificant relationships. This result indicates that the dependency of the austral-summer jet trend on model mean bias, which was evident in CCMVal2 simulations (Son *et al* 2010), is not clear in CCMI models, supporting the recent studies of Simpson and Polvani (2016) and Seviour *et al* (2017).

4. Summary and discussion

This study updates previous studies based on the CCMVal2 simulations by examining the state-of-the-art CCMs that participated in the CCMI project (Eyring *et al* 2013b). Most of these models are successors to the CCMVal2 models with improved chemistry (especially in the troposphere). Six models are also coupled with an ocean. Both CCMI-C1 and CCMI-C2 simulations, which differ mainly in their sea surface temperature and sea ice conditions, exhibit quantitatively similar multi-model mean trends over the period of 1960–2000 that are characterized by the poleward intensification of the austral-summer jet. The resulting trends are also quantitatively similar to the ones derived from the CCMVal2 and CMIP5 high-top models. This result suggests that Antarctic ozone-hole-induced tropospheric changes are not strongly sensitive to the specific chemistry-atmosphere-ocean coupling (Seviour *et al* 2017). The sensitivity of the austral-summer circulation changes to the details of stratospheric ozone forcing, reported in previous studies (Gillett *et al* 2009, Waugh *et al* 2009, Staten *et al* 2012, Neely *et al* 2014, Li *et al* 2016), appears to be smaller than the inter-model spread (or uncertainty

of the ozone-hole-induced tropospheric circulation change) and hence is not easily detectable.

All analyses shown in this study are based on only one ensemble member from each model. Although this allows a fair comparison among the models, it could make the result sensitive to the internal variability. To quantify the importance of internal variability, the analyses are repeated by considering all ensemble members. Here, multiple ensemble members (typically two or three) are available from six CCMI-C1 and seven CCMI-C2 models. Although not shown, overall results are not sensitive to the number of ensemble members. The multi-model mean trend based on only one ensemble member is quantitatively similar to the one derived from ensemble mean of each model.

Here we recall that all models analyzed in this study are forced not only by ozone depletion but also by all other external forcings such as increasing greenhouse gas concentrations and anthropogenic aerosol loadings. This implies that the austral-summer jet trends shown in this study are not solely driven by ozone depletion. Although it is well documented that ozone depletion is the major driver of historical SH-summer circulation change (e.g. Previdi and Polvani 2014), its relative importance against other forcings needs to be better quantified by examining the simulations with fixed ozone depleting substances (fODS) and fixed greenhouse gas (fGHG). Projected future circulation changes due to the anticipated ozone recovery also deserve further investigation.

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